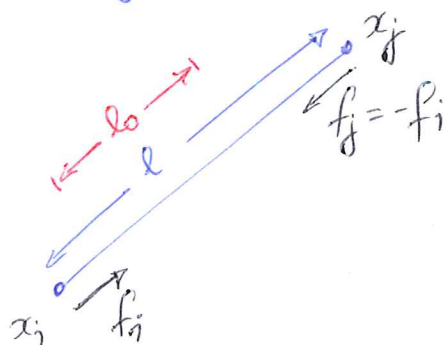


# Review

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\* Spring in n-dimensional space



$$\text{Force: } f_i = -k \left( \frac{l}{l_0} - 1 \right) \frac{x_j - x_i}{\|x_j - x_i\|}$$

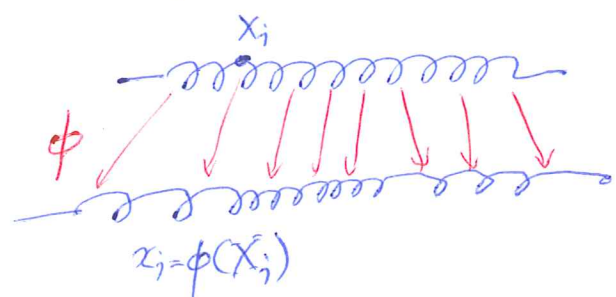
$$\text{Energy: } E = \underbrace{l_0}_{\text{Length}} \cdot \underbrace{\frac{k}{2} \left( \frac{l}{l_0} - 1 \right)^2}_{\psi\left(\frac{l}{l_0}\right) \Rightarrow \text{energy density}}$$

↓  
compression ratio

Relation of force-energy:

$$\vec{f}_i = - \frac{\partial E}{\partial \vec{x}_i} = \begin{pmatrix} -\partial E / \partial x_i \\ -\partial E / \partial y_i \\ -\partial E / \partial z_i \end{pmatrix}$$

The above works for a spring that is uniformly extended. If the extension is non-uniform, we consider the deformation function  $\phi(x)$ , which maps the "natural" (or "reference") position of each point, to its deformed location  $x = \phi(X)$



Example:

⇒ Spring moved along x-axis:

$$x = \phi(X) = X + c$$

⇒ Spring moved and twice extended

$$x = \phi(X) = 2X + c.$$

We associate the compression ratio  $(\frac{l}{l_0})$  with

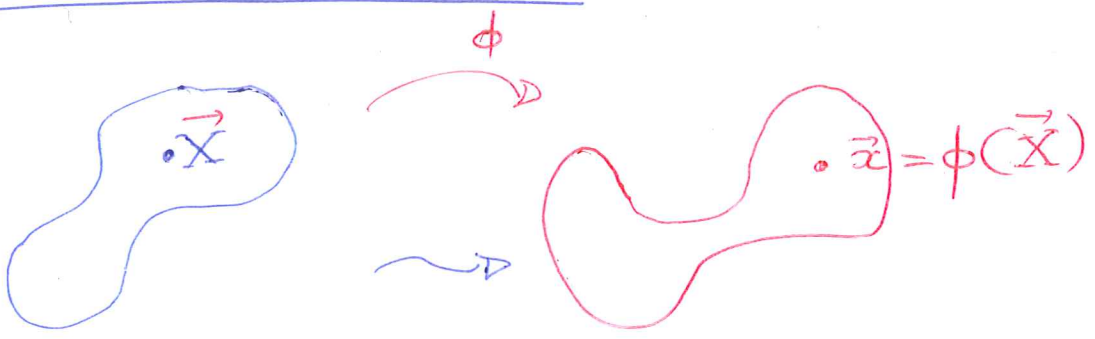
the derivative  $\phi'(X)$ : If the spring is uniformly extended,

then  $\phi'(X) = \text{const} = l/l_0$

Finally, energy  $E = \int_{X_L}^{X_R} \psi(\phi'(X)) dX$

( $X_L, X_R$  are the spring endpoints' reference locations)

Extension to volumetric solids



Reference shape

Deformed shape

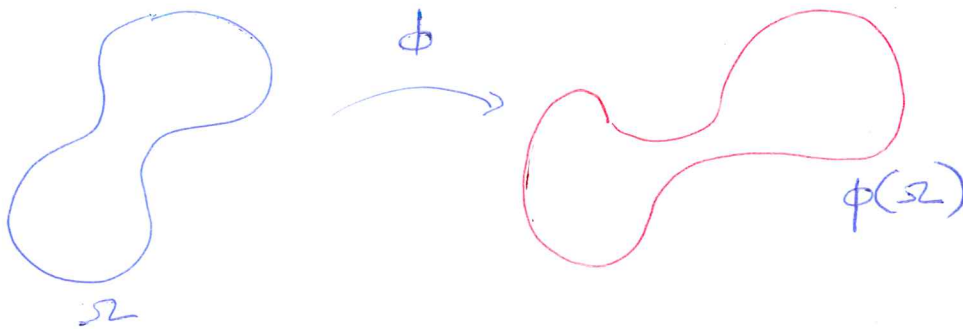
In this case  $\phi: \mathbb{R}^3 \rightarrow \mathbb{R}^3$   
 $\vec{X} \mapsto \vec{x} = \phi(\vec{X})$       $\phi \begin{pmatrix} X \\ Y \\ Z \end{pmatrix} = \begin{pmatrix} \phi_1(X, Y, Z) \\ \phi_2(X, Y, Z) \\ \phi_3(X, Y, Z) \end{pmatrix}$

The analogue of  $\phi'$  (as used in the spring energy) is the Jacobian matrix (called the "deformation gradient"  $F$ ).

$$F = \frac{d\vec{\phi}}{d\vec{X}} = \begin{pmatrix} \frac{\partial \phi_1}{\partial X} & \frac{\partial \phi_1}{\partial Y} & \frac{\partial \phi_1}{\partial Z} \\ \frac{\partial \phi_2}{\partial X} & \frac{\partial \phi_2}{\partial Y} & \frac{\partial \phi_2}{\partial Z} \\ \frac{\partial \phi_3}{\partial X} & \frac{\partial \phi_3}{\partial Y} & \frac{\partial \phi_3}{\partial Z} \end{pmatrix}$$

Similarly the energy stored in a deformed body is calculated as:

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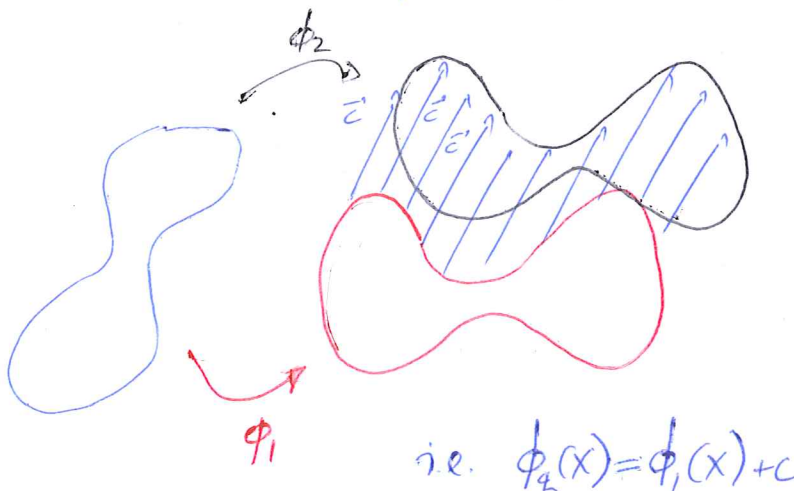
$$E = \int_{\Omega} \psi(F) dX$$

The exact definition of  $\psi(F)$  is a defining property of the material being modeled (e.g. steel, concrete, flesh, etc).

Properties:

→ Translational invariance

⇒ Consider 2 different deformations, which differ by a translational shift  $(\phi_1, \phi_2)$



Then:

$$\begin{aligned} F_1 &= \frac{\partial \phi_1}{\partial X} = \frac{\partial (\phi_2 + c)}{\partial X} \\ &= \frac{\partial \phi_2}{\partial X} = F_2 \end{aligned}$$

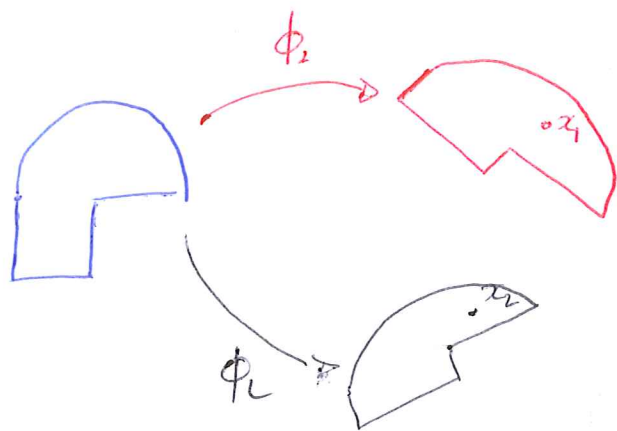
Thus, defining the energy density as a function of  $F$  automatically guarantees translational invariance, as:

$$F_1 = F_2 \rightsquigarrow \psi(F_1) = \psi(F_2)$$

$$\rightsquigarrow E[\phi_1] = \int \psi(F_1) = \int \psi(F_2) = E[\phi_2]$$

→ Rotational invariance

Consider 2 deformations that differ by a rotation, e.g.



$$\vec{x}_2 = R \vec{x}_1 + \vec{t}$$

$$\frac{d}{d\vec{x}} \downarrow$$

$$\boxed{F_2 = R F_1}$$

If the energy density satisfies  $\psi(F) = \psi(RF)$  for any rotation matrix  $R$ , then rotational invariance is automatically guaranteed

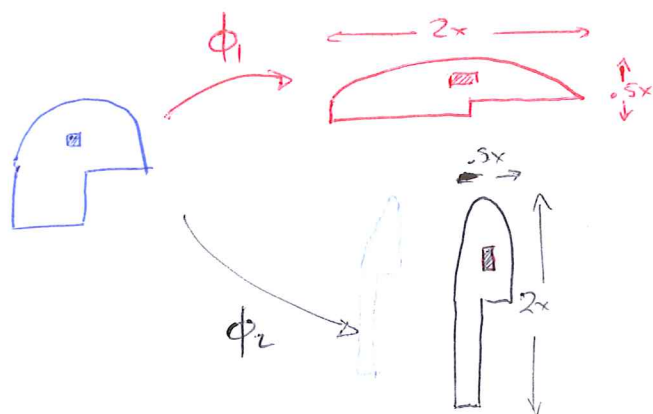
e.g.  $\psi(F) = \text{tr}(F^T F)$

$$\psi(RF) = \text{tr}(F^T \underbrace{R^T R}_{=I} F) = \text{tr}(F^T F) = \psi(F)$$

Rotational invariance is a desired property of  $\psi$  (although the cost of enforcing it may be non-trivial).

→ Isotropy

Consider 2 deformations where the same ratios of expansion/contraction are applied, but along different directions



Here we can prove that

$$F_1 = F_2 \cdot R$$

↑ post-multiply  
w/ rotation  
matrix

For certain materials we would expect the energy of these 2 deformations to be the same. Consider the deformation of an infinitesimal square section (mark by the shaded square) to see how this would be justified. This behavior is called isotropy and is characteristic of substances like rubber, jello, etc. Other materials such as woven fabrics, steel-reinforced concrete & muscles are not isotropic (there are directions along the material that resist deformation more than others).

If we want a material to be isotropic, we need to ensure that  $\psi(FR) = \psi(F)$  for all rotations  $R$ .

If we want to enforce both rotation invariance and isotropy, we need  $\psi(R_1 F R_2) = \psi(F)$  for any rotations  $R_1, R_2$ .

e.g.  $\psi = \det(F)$

$$\psi(R_1 F R_2) = \det(R_1 F R_2) = \overset{1}{\det(R_1)} \det(F) \overset{1}{\det(R_2)} = \psi(F)$$

We will see more examples of energy densities

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later, but here are two very simple academic examples:

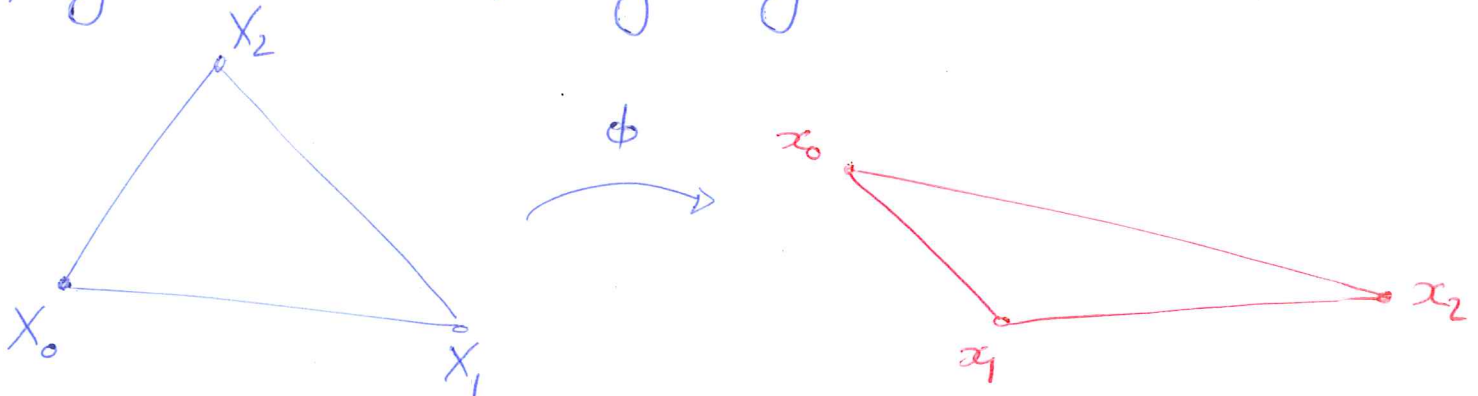
$$\Psi = \text{tr}(F^T F) = \|F\|_F^2 \quad \leadsto \text{"Zero rest-area membrane"} \\ \text{(Remember, } \|M\|_F^2 = \sum_{i,j=1}^n M_{ij}^2 \text{)}$$

We can also show that

$J := \det(F)$  is the % change in the volume of an infinitesimal region. An incompressible material satisfies  $J=1$

$$\Psi = c \cdot (J-1)^2 \quad \leadsto \text{A material that resists volume change (but nothing beyond that).}$$

Going back to a deforming triangle:



Here, we do not have access to the deformation function  $\phi(X)$ ; we only know how it transforms the vertices  $x_i = \phi(X_i)$ . In order to formulate an energy, we fit a model for  $\phi$  to these values, which is given by a linear relation

$$\vec{x} = \phi(\vec{X}) = A\vec{X} + \vec{t} \\ \begin{matrix} \uparrow & \uparrow \\ 2 \times 2 & 2\text{-vector} \end{matrix}$$

Observations :

- Such a model fit will be specific to the triangle  $t$  in question, i.e., we can write  $\phi^{(t)}(X) = A^{(t)} X + t$  per triangle
- Since  $F = \frac{\partial \phi}{\partial X} = \frac{\partial}{\partial X} (AX + t) = A^{=const}$ , the model really

takes the form  $\vec{x} = \phi(\vec{X}) = F \cdot \vec{X} + t$

(once again  $F$  is constant per triangle, but different from triangle to triangle).

- We can find  $F$ , as follows

$$\left. \begin{aligned} x_0 &= F X_0 + t \\ x_1 &= F X_1 + t \\ x_2 &= F X_2 + t \end{aligned} \right\} \Rightarrow \begin{aligned} x_1 - x_0 &= F (X_1 - X_0) \\ x_2 - x_0 &= F (X_2 - X_0) \end{aligned}$$

Finally we can write the last 2 equations as a single matrix equation:

$$\left[ \vec{x}_1 - \vec{x}_0 \mid \vec{x}_2 - \vec{x}_0 \right] = F \left[ \vec{X}_1 - \vec{X}_0 \mid \vec{X}_2 - \vec{X}_0 \right]$$

↖ ↗  
2 columns

or  $\begin{bmatrix} x_1 - x_0 & x_2 - x_0 \\ y_1 - y_0 & y_2 - y_0 \end{bmatrix} \stackrel{:= D_s}{=} F \begin{bmatrix} X_1 - X_0 & X_2 - X_0 \\ Y_1 - Y_0 & Y_2 - Y_0 \end{bmatrix} \stackrel{:= D_m}{=}$

$D_S$  = deformed shape matrix (variable overtime)

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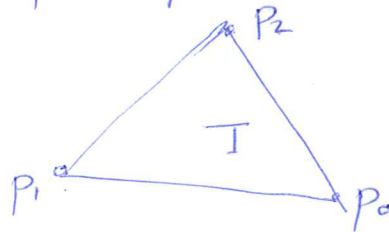
$D_m$  = reference shape matrix (constant)

since  $D_S = F D_m \Rightarrow \boxed{F = D_S D_m^{-1}}$

### Implementation detail

Since  $D_m$  is constant, we precompute & store  $D_m^{-1}$  ahead of time.

### Force definition process



Stage 1: Determine  $F = D_S D_m^{-1}$  from current positions.

$$(F = F(\vec{x}_0, \vec{x}_1, \vec{x}_2) = D_S(\vec{x}_0, \vec{x}_1, \vec{x}_2) \cdot D_m^{-1})$$

Stage 2: Plug into  $\psi(F)$

$$\psi(\vec{x}_0, \vec{x}_1, \vec{x}_2) = \psi(F(\vec{x}_0, \vec{x}_1, \vec{x}_2))$$

Stage 3: Energy of entire triangle  $T$

$$E^T = \int_T \psi(F) dX = \text{Vol}(T) \cdot \psi(F^T)$$

$T \rightarrow \text{constant on } T!$

$$= \text{Vol}(T) \cdot \psi(\vec{x}_0, \vec{x}_1, \vec{x}_2)$$

Stage 4: Compute forces on nodes as

$$\vec{f}_i = - \frac{\partial}{\partial \vec{x}_i} E(\vec{x}_0, \vec{x}_1, \vec{x}_2)$$



Doing all these steps explicitly would be tricky, since the final expression for  $E(x_0, x_1, \dots)$  can be elaborate and tricky to differentiate.

Fortunately, we can show that there is a compact expression for the forces, that results from this derivation. If we define the "force matrix"  $G$ , as follows:

$$G = \begin{bmatrix} \vec{f}_1 \\ \vdots \\ \vec{f}_2 \end{bmatrix} \text{ we can show that}$$

$$G = -\text{Vol}(T) \cdot P(F) \cdot D_m^{-T}$$

and the force on the last node is given by  $\vec{f}_0 = -\vec{f}_1 - \vec{f}_2$  (as forces need to balance out).

$P$  is called the 1st Piola-Kirchhoff stress matrix, and for every definition of  $\Psi(F)$  there is a corresponding definition of  $P(F)$ .

Note: It can be shown that  $\text{Vol}(T) = \begin{cases} \frac{1}{2} |\det D_m| & \text{in 2D} \\ \frac{1}{6} |\det D_m| & \text{in 3D} \end{cases}$

We typically precompute  $B_m = -\text{Vol}(T) D_m^{-T}$  ahead of time, and then calculate  $G = P \cdot B_m$ .

## Examples of material models:

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Note: For all of the following material models, we reference the 2 constants  $\mu$  &  $\lambda$ , called the Lamé parameters of the material. Their values are given as:

$$\mu = \frac{k}{2(1+\nu)} \quad \lambda = \frac{k\nu}{(1+\nu)(1-2\nu)}$$

where  $k$  = Young's modulus (a measure of stiffness)

$\nu$  = Poisson's ratio (ranging from  $0 \leq \nu < 0.5$  for most materials)

$\nu = 0$  Fully compressible (eg. cork)

$\nu = 0.5$  Fully incompressible (jello, rubber)

$\nu < 0$  exotic materials...

Remember: Frobenius norm of a matrix:

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$$\|M\|_F^2 = \sum_{ij} M_{ij}^2$$

## Linear elasticity

$$\Psi(F) = \mu \|\varepsilon\|_F^2 + \frac{\lambda}{2} \text{tr}(\varepsilon)^2$$

$$P(F) = 2\mu \varepsilon + \lambda \text{tr}(\varepsilon) \cdot I$$

where  $\varepsilon = \frac{1}{2}(F+F^T) - I$  "small strain tensor"

⇒ Not rotationally invariant  
(strictly speaking, not isotropic either.)

Benefit: Forces are a linear function of  $x_i$ 's!

$$D_S = [\bar{x}_1 - \bar{x}_0 \mid \bar{x}_2 - \bar{x}_0] \Rightarrow \text{linear on } x$$

$$F = D_S D_m^{-1} \Rightarrow \text{linear on } x$$

↑ const

$$P = \varepsilon = \frac{1}{2}(F+F^T) - I \Rightarrow \text{linear on } x$$

$$P = 2\mu \varepsilon + \lambda \text{tr}(\varepsilon) \cdot I \Rightarrow \text{linear on } x.$$

$$[f_1, f_2] = G = -\text{Vol}(CT) P D_m^{-1} \Rightarrow \text{linear on } x$$

Ultimately, we have 
$$\underline{f} = \begin{pmatrix} f_1 \\ f_2 \\ \vdots \\ f_n \end{pmatrix} = K \begin{pmatrix} x_1 \\ x_2 \\ \vdots \\ x_n \end{pmatrix} + b$$

↑ const.

which makes things easy with Backward Euler.

## St-Venant Kirchhoff material

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$$\Psi(F) = \mu \|E\|_F^2 + \frac{\lambda}{2} \text{tr}^2(E)$$

$$P(F) = F \left\{ 2\mu E + \lambda \text{tr}(E) \cdot I \right\}$$

where  $E = \frac{1}{2}(F^T F - I)$ . "Green strain tensor"

⇒ Isotropic, rotationally invariant

⇒  $f(x)$  is a nonlinear function

## Co-rotational (linear) elasticity

Let  $F = U \cdot \Sigma \cdot V^T$  be the singular value decomposition of  $F$ . Then  $F = U(V^T V) \Sigma V^T = \underbrace{(UV^T)}_R \underbrace{(V \Sigma V^T)}_S = \begin{matrix} R & S \\ \uparrow & \uparrow \\ \text{rotation} & \text{symmetric} \end{matrix}$

is called the polar decomposition of  $F$

$$\Psi(F) = \mu \|E^R\|_F^2 + \frac{\lambda}{2} \text{tr}^2(E^R)$$

$$P(F) = R \left\{ 2\mu E^R + \lambda \text{tr}(E^R) I \right\}$$

where  $E^R = S - I$  ( $S$ : from polar decomposition  $F = RS$ )

⇒ Isotropic, rotationally invariant

⇒ Easier to make work w/ Backward Euler than St-Venant-Kirchhoff.